

## SFHx

### EoS Submission Details

EoS name	SFHx
category	hadronic
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### Abstract

This is the SFHx EOS table [1] which is based on the statistical model with excluded volume and interactions of Hempel and Schaffner-Bielich (HS) [2] with RMF interactions SFHx [1]. Contributions of neutrons, anti-neutrons, protons, anti-protons, electrons, positrons, photons, and nuclei are included. For the masses of nuclei, FRDM [3] was used. The details of the underlying EOS model can be found in Ref. [2], where the TMA interactions were used. The manual from the web page

<http://phys-merger.physik.unibas.ch/~hempel/eos.html>

gives further information about the EOS table. On this web page, also routines are available which allow to determine the abundances of all nuclei for all conditions. Applications of HS EOS for various different RMF interactions in supernova simulations can be found in Refs. [2,4].

### References to the original work

1. M. Hempel and J. Schaffner-Bielich, Nucl. Phys. A **837** (2010) 210.
2. A.W. Steiner, M. Hempel, and T. Fischer (2012), arXiv:1207.2184.
3. P. Möller, J.R. Nix, and K.-L. Kratz, Atomic Data and Nuclear Data Tables **66** (1997) 131.

### Further References

4. M. Hempel, T. Fischer, J. Schaffner-Bielich, and M. Liebendörfer, Astrophys. J. **748** (2012) 70.

## Nuclear Matter Properties<sup>1</sup>

	Quantity	Unit	
$n_S$	saturation density in symmetric matter	$\text{fm}^{-3}$	0.1602
$E_0$	binding energy per baryon at saturation	MeV	16.16
$K$	incompressibility	MeV	238.8
$K'$	skewness	MeV	-457.2
$J$	symmetry energy	MeV	28.67
$L$	symmetry energy slope parameter	MeV	23.18
$K_{sym}$	symmetry incompressibility	MeV	-40.0

## Neutron Star Properties<sup>1</sup>

	Quantity	Unit	
$M_{max}$	maximum mass	$M_{\text{sun}}$	2.13
$M_{DU,e}$	mass at DUrca threshold (1/9) w/o $\mu^-$	$M_{\text{sun}}$	-
$R_{M_{max}}$	radius at maximum NS mass	km	10.77
$R_{1.4}$	radius at 1.4 $M_{\text{sun}}$ NS mass	km	12.0

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<sup>1</sup>0-values indicate, that the corresponding data is not provided.

## **eos.thermo**

eos.thermo and the three grid defining files are CompOSE standard data files and by definition available. eos.thermo does not necessarily provide all possible data.

table dimension                    3  
table type                         1  
total number of grid points   1584360

Range and density (#) of the grid parameters:

	Quantity	Unit	min	max	#
T	Temperature	MeV	0.10000000E+00	0.15848932E+03	81
$n_b$	Baryon Nr Density	$\text{fm}^{-3}$	0.10000000E-11	0.10000000E+02	326
$Y_q$	Charge Fraction		0.10000000E-01	0.60000000E+00	60

T,  $n_b$ , and  $Y_q$  are stored in eos.t, eos.nb, and eos.yq, respectively.

## **additional quantities in eos.thermo**

none defined

### Further Available Data Files

Files and quantities listed in the following are provided beyond CompOSE's core requirements as outlined in Sec.4.2. of the CompOSE manual.

**eos.compo** : available

index	particle
0	$e^-$
10	n
4002	${}^2_4\text{He}$
11	p
3002	${}^2_3\text{He}$
3001	${}^1_3\text{H}$
2001	${}^1_2\text{H}$
	- end of table -

The listed particle number fractions of electrons, protons, and neutrons are net fractions, i.e., they are given by the difference between the corresponding particle and anti-particle number density.

Further particle sets are defined. One set of quadruples for an average “heavy” nucleus, see Table 7.2 of the manual.

index	description
999	group of all other considered nuclei which are not listed above (averaged)
	- end of table -

**eos.micro** : available

index	quantity	particle
10041	Dirac effective mass divided by particle mass $m_i^D/m_i$	n
11041	Dirac effective mass divided by particle mass $m_i^D/m_i$	p
	- end of table -	

## Description of Phases

Fill this part briefly, in particular if several phases occur. In this latter case characterize the transition(s).

**PHASE INDEX #1:**

NSE phase, i.e., a mixture of nuclei and nucleons

**PHASE INDEX #3:**

pure RMF, i.e., only nucleons

**PHASE INDEX #2:**

Maxwell transition region between phase 1 and 3, assuming local charge neutrality and locally fixed  $Y_e$ .