## QHC21T $A_T$

### **EoS Submission Details**

EoS name	QHC21T $A_T$
category	hybrid
submitted by	Toru Kojo
affiliation	Central China Normal University
e-mail contact	kojo.toru@gmail.com
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### Abstract

This table corresponds to the zero temperature and  $\beta$ -equilibrium unified EoS by Kojo et al. [1]. The EoS are divided into four distinct domains: the crust, nuclear liquid, hadron-quark crossover, and quark matter domains. For each domain we assign an equation of state as

Crust to nuclear liquid	Togashi [2]	$[10^{-9}n_0 \le n_B \le 1.5n_0]$
Crossover	QHC21T $A_T$ [1]	$[1.5n_0 \le n_B \le 3.5n_0]$
Quark matter	QHC21T $A_T$ [1]	$[3.5n_0 \le n_B \le 10n_0]$

The quark matter EoS (including up-, down-, and strange-quarks) in the QHC21T was calculated using the NJL model within the mean field approximation. As variable parameters, we choose  $g_V$  and H which quantify the strength of the repulsive densitydensity interaction and the attractive paring-interaction between quarks, respectively. For QHC21T A<sub>T</sub>, we consider  $(g_V, H)/G = (0.90, 1.50)$  [A<sub>T</sub>] which are compatible with the hadron physics. The QHC21T satisfies the empirical constraints from neutron stars as well as the causality and thermodynamic consistency.

### References to the original work

- T. Kojo, G. Baym, and T. Hatsuda, "QHC21 equation of state of neutron star matter - in light of 2021 NICER data," arXiv: 2111.11919, https://arxiv.org/pdf/2111.11919.pdf
- H. Togashi, K. Nakazato, Y. Takehara, S. Yamamuro, H. Suzuki, and M. Takano, *"Nuclear equation of state for core-collapse supernova simulations with realistic nuclear forces,"* Nucl. Phys. A 961, 78 (2017), https://doi.org/10.1016/j.nuclphysa.2017.02.010

## Nuclear Matter Properties<sup>1</sup>

	Quantity	Unit		
$n_S$	saturation density in symmetric matter	$\mathrm{fm}^{-3}$	0.16	
$E_0$	binding energy per baryon at saturation	MeV	16.1	
K	incompressibility	MeV	245.0	
K'	skewness	$\mathrm{MeV}$	0	
J	symmetry energy	$\mathrm{MeV}$	29.1	
L	symmetry energy slope parameter	$\mathrm{MeV}$	38.7	
$K_{sym}$	symmetry incompressibility	$\mathrm{MeV}$	0	
$U_{\Lambda}$	$\Lambda$ -potential at saturation	$\mathrm{MeV}$	0	
$U_{\Sigma}$	$\Sigma$ -potential at saturation	$\mathrm{MeV}$	0	
$U_{\Xi}$	$\Xi$ -potential at saturation	MeV	0	

# Neutron Star Properties<sup>1</sup>

	Quantity	Unit	
M <sub>max</sub>	maximum mass	$M_{sun}$	2.13
$M_{DU,e}$	mass at DUrca threshold (1/9) w/o $\mu^-$	$M_{sun}$	0
$R_{M_{max}}$	radius at maximum NS mass	$\mathrm{km}$	10.9
$R_{1.4}$	radius at 1.4 $M_{sun}$ NS mass	$\mathrm{km}$	11.8
$ ilde{\Lambda}$	tidal deformability for GW170817 at a mass ratio of $q = 0.8$		0

### eos.thermo

eos.thermo and the three grid defining files are CompOSE standard data files and by definition available.

table dimension	1
table type	1
total number of grid points	507

<sup>&</sup>lt;sup>1</sup>0-values indicate, that the corresponding data is not provided.

Range and density (#) of the grid parameters:

	Quantity	Unit	$\min$	$\max$	#	
Т	Temperature	MeV	0	0	1	
$\mathbf{n}_b$	Baryon Nr Density	${\rm fm}^{-3}$	7.59E-011	1.64	507	
$\mathbf{Y}_q$	Charge Fraction		0	0	1	

T,  $\mathbf{n}_b,$  and  $\mathbf{Y}_q$  are stored in eos.t, eos.nb, and eos.yq, respectively.

### **Further Available Data Files**

Files and quantities listed in the following are provided beyond CompOSE's core requirements as outlined in Sec.4.2. of the CompOSE manual.

 $\mathbf{e}$ os.compo : not available

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