**R(DD2Y**
$$\Delta$$
)  $x_{\sigma\Delta} = 1.2$ ;  $x_{\omega\Delta} = 1.3$ ;  $x_{\rho\Delta} = 1.0$ ;

#### **EoS Submission Details**

EoS name  $R(DD2Y\Delta) x_{\sigma\Delta} = 1.2; x_{\omega\Delta} = 1.3; x_{\rho\Delta} = 1.0;$ 

category Hadronic

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sheet creation date January 23, 2022

#### **Abstract**

This hadronic EOS table accounts for hyperons and  $\Delta(1232)$  resonances in addition to nucleons [1,2]. The nucleonic RMF effective interaction is DD2 [3]. The coupling constants of exotic species to different mesonic fields are provided in the table nearby. For the clusterized phase occurring at sub-saturation densities we use data in HS(DD2) available on CompOSE; they have been obtained within a statistical model with excluded volume and interactions [4]. The transition from unhomogeneous matter to homogeneous matter is done by minimizing the free energy density. For the masses of nuclei, FRDM [5] was used. Contributions of electrons, positrons and photons are included.

#### References to the original work

- 1. Ad. R. Raduta et al., in preparation (2022).
- 2. Ad. R. Raduta, M. Oertel, A. Sedrakian, MNRAS 499 (2020) 914-931.
- 3. S. Typel, G. Ropke, T. Klahn, D. Blaschke, and H.H. Wolter, Phys. Rev. C 81 (2010) 015803.
- 4. M. Hempel and J. Schaffner-Bielich, Nucl. Phys. A 837 (2010) 210.
- 5. P. Moller, J.R. Nix, and K.-L. Kratz, Atomic Data and Nuclear Data Tables 66 (1997) 131.

#### Coupling constants of exotic species to meson fields

expressed in terms of the coupling constants of the nucleons N to the meson fields,  $x_{mB} = g_{mB}/g_{mN}$ .

```
coupling constant
                                           value
                                           0.6154
               x_{\sigma\Lambda}
                                           0.3259
               x_{\sigma\Xi}
                                           0.4740
               x_{\sigma\Sigma}
                                           1.2000
               x_{\sigma\Delta}
               x_{\omega\Lambda}
                                           2/3
                                           1/3
               x_{\omega\Xi}
                                           2/3
               x_{\omega\Sigma}
                                           1.3000
               x_{\omega\Delta}
                                           0
               x_{\rho\Lambda}
              x_{\rho\Xi}
                                           1
                                           2
               x_{\rho\Sigma}
                                           1
               x_{\rho\Delta}
                                           \begin{array}{c} -\sqrt{2}/3 \\ -\sqrt{2}/3 \end{array}
               x_{\phi\Lambda}
               x_{\phi\Xi}
                                           -2\sqrt{2}/3
               x_{\phi\Sigma}
                                           0
               x_{\phi\Delta}
                                           - end of table -
```

## **Nuclear Matter Properties**<sup>1</sup>

	Quantity	$\operatorname{Unit}$	
$\overline{n_S}$	saturation density in symmetric matter	$\mathrm{fm}^{-3}$	0.1491
$E_0$	binding energy per baryon at saturation	MeV	16.02
K	incompressibility	MeV	242.7
K'	skewness	MeV	168.7
J	symmetry energy	MeV	31.67
L	symmetry energy slope parameter	MeV	55.03
$K_{sym}$	symmetry incompressibility	MeV	-93.23

# Neutron Star Properties<sup>1</sup>

	Quantity	Unit	
$\overline{M_{max}}$	maximum mass	$M_{sun}$	2.032
$M_{DU,e}$	mass at DUrca threshold (1/9) w/o $\mu^-$	$M_{sun}$	1.60
$R_{M_{max}}$	radius at maximum NS mass	$\mathrm{km}$	11.44
$R_{1.4} \  ilde{\Lambda}$	radius at 1.4 $M_{sun}$ NS mass	$\mathrm{km}$	13.25
$ ilde{\Lambda}$	tidal deformability for GW170817 at a mass ratio of $q=0.8$		770

### eos.thermo

eos. thermo and the three grid defining files are CompOSE standard data files and by definition available.

# additional quantities in eos.thermo

none defined

 $\begin{array}{ccc} \text{table dimension} & & 3 \\ \text{table type} & & 1 \\ \text{total number of grid points} & 1377120 \end{array}$ 

<sup>&</sup>lt;sup>1</sup>0-values indicate, that the corresponding data is not provided.

Range and density (#) of the grid parameters:

	Quantity	Unit	min	max	#	
$\overline{\mathrm{T}}$	Temperature	MeV	$0.1~\mathrm{MeV}$	100 MeV	76	
$n_b$	Baryon Nr Density	$\rm fm^{-3}$	$10^{-12} \text{ fm}^{-3}$	$1.0964782~{\rm fm}^{-3}$	302	
$Y_q$	Charge Fraction		0.01	0.6	60	

T,  $\mathbf{n}_b,$  and  $\mathbf{Y}_q$  are stored in eos.t, eos.nb, and eos.yq, respectively.

#### **Further Available Data Files**

Files and quantities listed in the following are provided beyond CompOSE's core requirements as outlined in Sec.4.2. of the CompOSE manual.

### eos.compo: available

$$\begin{array}{c|cccc} \text{index} & \text{particle} \\ 0 & \text{e} \\ 10 & \text{n} \\ 11 & \text{p} \\ 20 & \Delta^- \\ 21 & \Delta^0 \\ 22 & \Delta^+ \\ 23 & \Delta^{++} \\ 100 & \Lambda \\ 110 & \Sigma^- \\ 111 & \Sigma^0 \\ 112 & \Sigma^+ \\ 120 & \Xi^- \\ 121 & \Xi^0 \\ 4002 & {}^4_2\text{He} \\ 3002 & {}^3_2\text{He} \\ 3001 & {}^3_1\text{H} \\ 2001 & {}^2_1\text{H} \\ & \text{- end of table -} \\ \end{array}$$

further particle sets are defined. One set of quadruples for an average "heavy" nucleus, see Table 7.2 of the manual.

eos.micro: available

index	quantity	particle
10041	Dirac effective mass divided by particle mass $m_i^D/m_i$	n
11041	Dirac effective mass divided by particle mass $m_i^D/m_i$	p
20041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Delta^-$
21041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Delta^0$
22041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Delta^+$
23041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Delta^{++}$
100041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Lambda$
110041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Sigma^-$
111041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Sigma^0$
112041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Sigma^+$
120041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Xi^-$
121041	Dirac effective mass divided by particle mass $m_i^D/m_i$	$\Xi^0$
	- end of table -	

## **Description of Phases**

Fill this part briefly, in particular if several phases occur. In this latter case characterize the transition(s).

```
PHASE INDEX #1:
```

NSE phase, i.e., a mixture of nuclei and baryons

PHASE INDEX #3:

homogeneous matter

PHASE INDEX #2:

Maxwell transition region between phase 1 and 3, assuming local charge neutrality and locally fixed  $Y_e$  .