

# FSU2H

## EoS Submission Details

EoS name	FSU2H
category	Inner crust
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## Abstract

This inner crust EoS, including nonspherical pasta phases, was calculated within a self-consistent Thomas-Fermi approach [1] for  $\beta$ -equilibrium matter at zero temperature. The EoS parameterisation used was the non-linear RMF FSU2H model [2]. This inner crust EoS has been published in Ref. [3]. Matched to model BsK22 [4] for the description of the outer crust below  $n_B = 2 \times 10^{-3} \text{ fm}^{-3}$ .

## References to the original work

1. F. Grill. H. Pais, C. Providência, I. Vidaña, and S. Avancini, Phys. Rev. C **90**, 045803 (2014).
2. R. Negreiros, L. Tolos, M. Centelles, A. Ramos, and V. Dexheimer, Astrophys. J. **863**, 104 (2018).
3. C. Providência, M. Fortin, H. Pais, and A. Rabhi, Front. Astron. Space Sci. **6**, 13 (2019).
4. J. M. Pearson, N. Chamel, A. Y. Potekhin, A. F. Fantina, C. Ducoin, A. K. Dutta, and S. Goriely, MNRAS **481**, 2994 (2018); <https://compose.obspm.fr/eos/210>

## Nuclear Matter Properties<sup>1</sup>

	Quantity	Unit	
$n_S$	saturation density in symmetric matter	$\text{fm}^{-3}$	0.1505
$E_0$	binding energy per baryon at saturation	MeV	-16.28
$K$	incompressibility	MeV	238
$K'$	skewness	MeV	-24.6
$J$	symmetry energy	MeV	30.5
$L$	symmetry energy slope parameter	MeV	44.5
$K_{sym}$	symmetry incompressibility	MeV	87

## Neutron Star Properties<sup>1</sup>

	Quantity	Unit	
$M_{max}$	maximum mass	$M_{\text{sun}}$	2.37
$M_{DU,e}$	mass at DUrca threshold (1/9) w/o $\mu^-$	$M_{\text{sun}}$	0
$R_{M_{max}}$	radius at maximum NS mass	km	12.43
$R_{1.4}$	radius at 1.4 $M_{\text{sun}}$ NS mass	km	13.29
$\tilde{\Lambda}$	tidal deformability for GW170817 at a mass ratio of $q = 0.8$	-	856.27

## eos.thermo

eos.thermo and the three grid defining files are CompOSE standard data files and by definition available. In eos.thermo, an extra quantity is added (last column). It refers to the geometry of the correspondent pasta phase, represented by an integer, with 0 for the outer crust, and 1 - droplets, 2 - rods, 3 - slabs, 4- tubes, 5- bubbles, and 6 corresponds to the core (homogeneous matter).

table dimension                    1  
table type                            1  
total number of grid points    300

Range and density (#) of the grid parameters:

	Quantity	Unit	min	max	#
T	Temperature	MeV	0	0	1
$n_b$	Baryon Nr Density	$\text{fm}^{-3}$	$0.467962 \times 10^{-9}$	$0.87 \times 10^{-1}$	300
$Y_q$	Charge Fraction		0	0	1

T,  $n_b$ , and  $Y_q$  are stored in eos.t, eos.nb, and eos.yq, respectively.

<sup>1</sup>0-values indicate that the corresponding data is not provided.

### Further Available Data Files

Files and quantities listed in the following are provided beyond CompOSE's core requirements as outlined in Sec.4.2. of the CompOSE manual.

**eos.mr** : This file provides the gravitational mass (in solar masses), the radius (in km), and the tidal deformability of a family of stars computed for this unified inner-crust-core RMF EoS model, with the BsK22 outer crust.

**eos.compo** : available

4 particle pairs (neutrons, protons, electrons, muons) and one quadruple for heavy nucleus.

Phase index # 0: outer crust

Phase index # 7: inner crust

Phase index # 6: homogeneous matter

**eos.micro** : not available