#### **EoS Submission Details**

EoS name	SPG(M4)
category	Inner crust
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## Abstract

This crust EoS for  $\beta$ -equilibrium matter at zero temperature is constructed with a BSk22 outer crust [1] below  $n_B = 2 \times 10^{-3}$  fm<sup>-3</sup> and inner crust EoS calculated within a Compressible Liquid Drop approach [2] up to  $n_B = 0.98 \times 10^{-1}$  fm<sup>-3</sup>. The EoS parameterisation used was one of the density-dependent RMF models with the GDFM functional for the couplings [3] generated in a Bayesian analysis and that agrees with both astrophysical and nuclear constraints [4].

### References to the original work

- J. M. Pearson, N. Chamel, A. Y. Potekhin, A. F. Fantina, C. Ducoin, A. K. Dutta, and S. Goriely, MNRAS 481, 2994 (2018); https://compose.obspm.fr/eos/210.
- 2. G. Baym, H. A. Bethe and C. J. Pethick, Nuclear Physics A 175, 225 (1971).
- P. Gogelein, E. N. E. van Dalen, C. Fuchs and H. Muther, Phys. Rev. C 77, 025802 (2008).
- 4. L. Scurto, H. Pais and F. Gulminelli in preparation

## Nuclear Matter Properties<sup>1</sup>

	Quantity	Unit		
$n_S$	saturation density in symmetric matter	$\mathrm{fm}^{-3}$	0.162	
$E_0$	binding energy per baryon at saturation	MeV	-15.8	
K	incompressibility	MeV	246	
K'	skewness	MeV	293	
J	symmetry energy	MeV	30.9	
L	symmetry energy slope parameter	MeV	40	
$K_{sym}$	symmetry incompressibility	MeV	-111.8	

# Neutron Star Properties<sup>1</sup>

	Quantity	Unit	
$M_{max}$	maximum mass	$M_{sun}$	2.35
$M_{DU}$	mass at DUrca threshold with $\mu^-$	$M_{\mathrm{sun}}$	Х
$R_{M_{max}}$	radius at maximum NS mass	$\mathrm{km}$	11.26
$R_{1.4}$	radius at 1.4 $M_{sun}$ NS mass	$\mathrm{km}$	12.31
$ ilde{\Lambda}$	tidal deformability for GW170817 at a mass ratio of $q = 0.8$	-	757.32

## eos.thermo

eos.thermo and the three grid defining files are CompOSE standard data files and by definition available.

table dimension1table type1total number of grid points375

Range and density (#) of the grid parameters:

	Quantity	Unit	$\min$	max	#	
Т	Temperature	MeV	0	0	1	
$\mathbf{n}_b$	Baryon Nr Density	${\rm fm}^{-3}$	$0.467962 \times 10^{-9}$	$0.9772 \times 10^{-1}$	375	
$\mathbf{Y}_q$	Charge Fraction		0	0	1	

T,  $n_b$ , and  $Y_q$  are stored in eos.t, eos.nb, and eos.yq, respectively.

<sup>&</sup>lt;sup>1</sup>0-values indicate that the corresponding data is not provided. X indicates DUrca threshold is not reached before the maximum mass.

## **Further Available Data Files**

Files and quantities listed in the following are provided beyond CompOSE's core requirements as outlined in Sec.4.2. of the CompOSE manual.

**eos.mr** : This file provides the radius (in km), the gravitational mass (in solar masses), the adimensional tidal deformability and the central density (in  $fm^{-3}$ ) of a family of stars computed for this unified inner-crust–core RMF EoS model, with the BSk22 outer crust.

eos.compo : available

4 particle pairs (neutrons, protons, electrons, muons) and one quadruple for heavy nucleus.

Phase index # 0: outer crust Phase index # 7: inner crust

eos.micro : not available