

APR(APR) with unified crust

EoS Submission Details

EoS name	APR(APR) with unified crust
category	Hadronic
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Abstract

This table represents the zero temperature and β -equilibrium EoS by Akmal, Pandharipande and Ravenhall using variational techniques [1], interaction $A18 + \delta v + UIX^*$. The transition to the pion condensed phase at high densities is modelled via an inhomogeneous mixed phase with a Gibbs construction, see chapter 5.12 of [2]. The crust has been reconstructed in a thermodynamically consistent and unified way using the CUTER tool [3,4], see [4] for details.

References to the original work

1. A. Akmal, V.R. Pandharipande and D.G. Ravenhall, Phys. Rev. C 58 (1998) 1804.
2. P. Haensel, A.Y. Potekhin, D.G. Yakovlev, Neutron Stars 1: Equation of State and Structure, Springer, New York (2007).
3. P.J. Davis, *et al*, Astron. Astrophys. 687, 44 (2024).
4. P.J. Davis *et al*, Eur. Phys. J. A (2025).

Nuclear Matter Properties¹

	Quantity	Unit	
n_S	saturation density in symmetric matter	fm^{-3}	0.16
E_0	binding energy per baryon at saturation	MeV	16.0
K	incompressibility	MeV	266.0
K'	skewness	MeV	–
J	symmetry energy	MeV	32.6
L	symmetry energy slope parameter	MeV	57.6
K_{sym}	symmetry incompressibility	MeV	–

Neutron Star Properties¹

	Quantity	Unit	
M_{max}	maximum mass	M_{sun}	2.19
$M_{DU,e}$	mass at DUrca threshold (1/9) w/o μ^-	M_{sun}	–
$R_{M_{max}}$	radius at maximum NS mass	km	9.94
$R_{1.4}$	radius at 1.4 M_{sun} NS mass	km	11.37
$\tilde{\Lambda}$	tidal deformability for a NS with $M = 1.4M_{sun}$		246

eos.thermo

eos.thermo and the three grid defining files are CompOSE standard data files and by definition available.

table dimension	1
table type	1
total number of grid points	1833

¹0-values indicate, that the corresponding data is not provided.

Range and density (#) of the grid parameters:

	Quantity	Unit	min	max	#
T	Temperature	MeV	0	0	1
n_b	Baryon Nr Density	fm^{-3}	1.e-11	1.34	1833
Y_q	Charge Fraction		0	0	1

T, n_b , and Y_q are stored in eos.t, eos.nb, and eos.yq, respectively.