

DD2F-SF quark-hadron model RDF 1.9 with electrons

EoS Submission Details

| | |
|---------------------|---|
| EoS name | DD2F-SF quark-hadron model RDF 1.9 with electrons |
| category | hybrid |
| submitted by | Niels-Uwe Friedrich Bastian |
| affiliation | University of Wroclaw |
| e-mail contact | niels-uwe.bastian(at)uwr.edu.pl |
| sheet creation date | March 15, 2021 |

Abstract

RDF 1.9 is from a set of equation of state models with a first-order phase transition from hadron to quark matter presented in Ref. [1]. This publication focuses on the super-saturated regime and for the treatment of nuclear clusters below saturation on the nuclear statistical equilibrium with excluded volume approach of Hempel and Schaffner-Bielich [2]. The relativistic density function approach, which is used for the high density regime and in particular the string-flip model for quarks was introduced in Ref. [3] with applications to neutron star configurations. The current extension of this model to finite temperature and arbitrary chargefractions was successfully applied to core-collapse supernova simulations [4] and binary neutron star merger simulations [5-7].

On the website of the author¹ you will find further information about the provided data, and the data in different formats. Please feel free to contact the author, if you run into problems with these tabulations.

¹<https://eos.bastian.science>

References to the original work

1. N.-U. F. Bastian, Phys. Rev. D 103.2, p. 023001 (2021).
doi: 10.1103/PhysRevD.103.023001.
2. M. Hempel and J. Schaffner-Bielich, Nucl. Phys. A 837 (2010) 210.
doi: 10.1016/j.nuclphysa.2010.02.010

Further References

3. M. A. R. Kaltenborn, N.-U. F. Bastian, and D. B. Blaschke, Phys. Rev. D 96, 056024.
doi: 10.1103/PhysRevD.96.056024.
4. T. Fischer, N.-U. F. Bastian, M.-R. Wu, P. Baklanov, E. Sorokina, S. Blinnikov, S. Typel, T. Klähn, D. B. Blaschke, Nature Astronomy 2, 980–986 (2018),
doi: 10.1038/s41550-018-0583-0
5. Andreas Bauswein, Niels-Uwe F. Bastian, David B. Blaschke, Katerina Chatziioannou, James A. Clark, Tobias Fischer, and Micaela Oertel, Phys. Rev. Lett. 122, 061102 (2019),
doi: 10.1103/PhysRevLett.122.061102
6. Andreas Bauswein, Sebastian Blacker, Vimal Vijayan, Nikolaos Stergioulas, Katerina Chatziioannou, James A. Clark, Niels-Uwe F. Bastian, David B. Blaschke, Mateusz Cierniak, and Tobias Fischer, Phys. Rev. Lett. 125, 141103 (2020),
doi: 10.1103/PhysRevLett.125.141103
7. Sebastian Blacker, Niels-Uwe F. Bastian, Andreas Bauswein, David B. Blaschke, Tobias Fischer, Micaela Oertel, Theodoros Soutanis, and Stefan Typel, Phys. Rev. D 102, 123023 (2020),
doi: 10.1103/PhysRevD.102.123023

Nuclear Matter Properties²

| | Quantity | Unit | |
|-----------|---|------------------|--------|
| n_S | saturation density in symmetric matter | fm^{-3} | 0.149 |
| E_0 | binding energy per baryon at saturation | MeV | 16.02 |
| K | incompressibility | MeV | 242.7 |
| K' | skewness | MeV | 168.8 |
| J | symmetry energy | MeV | 31.67 |
| L | symmetry energy slope parameter | MeV | 55.04 |
| K_{sym} | symmetry incompressibility | MeV | -93.23 |

Neutron Star Properties²

| | Quantity | Unit | |
|---------------|---|-----------|------|
| M_{max} | maximum mass | M_{sun} | 2.17 |
| $M_{DU,e}$ | mass at DUrca threshold (1/9) w/o μ^- | M_{sun} | N/A |
| $R_{M_{max}}$ | radius at maximum NS mass | km | N/A |
| $R_{1.4}$ | radius at 1.4 M_{sun} NS mass | km | N/A |

eos.thermo

eos.thermo and the three grid defining files are ComPOSE standard data files and by definition available. eos.thermo does not necessarily provide all possible data.

```
table dimension          3
table type               1
total number of grid points 1099800
```

Range and density (#) of the grid parameters:

| | Quantity | Unit | min | max | # |
|-------|-------------------|------------------|----------------|----------------|-----|
| T | Temperature | MeV | 0.1E+00 | 0.1202264E+03 | 81 |
| n_b | Baryon Nr Density | fm^{-3} | 0.1E-11 | 0.1E+02 | 235 |
| Y_q | Charge Fraction | | 0.10000000E-01 | 0.60000000E+00 | 60 |

T, n_b , and Y_q are stored in eos.t, eos.nb, and eos.yq, respectively.

²0-values indicate, that the corresponding data is not provided.

Further Available Data Files

Files and quantities listed in the following are provided beyond CompOSE's core requirements as outlined in Sec.4.2. of the CompOSE manual.

eos.compo : available

| index | particle |
|-------|-------------------|
| 0 | e^- |
| 10 | n |
| 11 | p |
| 4002 | ${}^2_4\text{He}$ |
| 3002 | ${}^2_3\text{He}$ |
| 3001 | ${}^1_3\text{H}$ |
| 2001 | ${}^1_2\text{H}$ |
| 500 | u |
| 501 | d |
| | - end of table - |

The listed particle number fractions are net fractions, i.e., they are given by the difference between the corresponding particle and anti-particle fractions. Further particle sets are defined.

| index | description |
|-------|--|
| 999 | Average fraction, mass and proton number for all nuclei not listed above |
| | - end of table - |

eos.micro : available

| index | quantity | particle |
|--------|---|----------|
| 10041 | Dirac effective mass divided by particle mass m_i^D/m_i | n |
| 11041 | Dirac effective mass divided by particle mass m_i^D/m_i | p |
| 500041 | Dirac effective mass divided by particle mass m_i^D/m_i | u |
| 501041 | Dirac effective mass divided by particle mass m_i^D/m_i | d |
| 10051 | relativistic vector self-energy V_i | n |
| 11051 | relativistic vector self-energy V_i | p |
| 500051 | relativistic vector self-energy V_i | u |
| 501051 | relativistic vector self-energy V_i | d |
| | - end of table - | |